Synthesis, Characterization, and Chemical Vapor Deposition Properties of Primary and Secondary Neopentylstibine. New Antimony Precursors for MOCVD

D. **Greg Hendershot,*J John C. Pazik, and Alan D. Berry**

Naval Research Laboratory, Materials Chemistry Branch, Code *6120,* Washington, D.C. *20375-5000*

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The primary and secondary neopentylstibines (neopentyl = $Me₃CCH₂$ -) have been synthesized by the reduction of the corresponding dihalo- and monohaloantimony compounds, respectively. Neopentylstibine was prepared by the reduction of $(Me_3CCH_2)SbBr_2$ with lithium aluminum hydride (LAH). The primary stibine is highly air-sensitive and exhibits a vapor pressure of 5.5 Torr at 0 °C. Dineopentylstibine was synthesized by LAH reduction of $(Me_3CCH_2)_2$ SbI. The secondary stibine is less sensitive to air than its primary congener but nonetheless decomposes when exposed to air. Dineopentylstibine is substantially less volatile than $(Me_3CCH_2)SbH_2$, exhibiting a vapor pressure of 0.5 Torr at 24 C. The detailed synthesis, spectroscopic characterizations, physical properties, and preliminary chemical vapor deposition experiments of

Introduction

The production of semiconductor films by chemical vapor deposition techniques is an extremely active research
area.² An ever-increasing demand for semiconductor An ever-increasing demand for semiconductor materials specifically tailored to meet desired requirementa fuels the efforts to understand materials production and fabrication. One of the most intriguing areas in semiconductors today is the antimonides, which find a wide variety of applications in infrared detectors, high-speed digital circuits, and quantum well structures.³

Although there is an increasing amount of ongoing work in antimonide-based semiconductors, these materials are less studied than other III-V systems such as InP and **GaAs.** This stems mostly from the difficulty associated with growing reproducible, high-quality antimonide films by metal organic chemical vapor deposition (MOCVD). There are at least two factors that affect the growth of antimonide-based thin films. First is the limited number of antimony sources currently available for use in CVD experiments. At this time, trimethylstibine $(M_{\theta_3}S_b)$ is the precursor of choice since it is the only volatile antimony source that is commercially available. Examination of the literature reveals that stibine $(SbH₃)$,⁴ triethylstibine $(Et₃Sb)⁵$ and recently triisopropylstibine $(i-Pr₃Sb)⁶$ have been explored **as** alternative antimony sources. Although these four sources have been used for the growth of acceptable films, each suffers from a drawback such as in-

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stability, low volatility, or high decomposition temperature. A second factor associated with antimonide film growth involves the relationship between the V/III ratio and the deposition temperature. Unlike the liberal growth conditions employed for **GaAs** that utilize a broad range of values for V/III ratios (always with excess AsH_3) over a wide temperature range, growth of antimonide films is more sensitive to the $\overline{V}/\overline{II}$ ratio employed at a specific deposition temperature.⁷ A slight variation of either A slight variation of either parameter from its optimum value often results in either low growth rates or poor morphology and composition.

In view of the success of t -BuAsH₂⁸ and t -BuPH₂⁹ in the growth of **GaAs** and InP, it is surprising that primary $(RSbH₂)$ and/or secondary stibines $(R₂SbH)$ have not been examined as alternative precursors to Me₃Sb. Mechanistic studies have shown that hydrogen atoms bound to the group V atom often reduce carbon contamination in film **growth?b** We believe the reasons organoantimony hydrides have not been considered as MOCVD precursors are related to the few known compounds of this type and to their lack of stability or volatility. To our knowledge, only two primary stibines and five secondary stibines have been reported in the literature to date. The methylstibines $Me₂SbH$ and $MeSbH₂$ are volatile liquids but reportedly decompose at room temperature and above -78 °C , respectively.¹⁰ Of the two remaining alkylstibines. (*t*-Of the two remaining alkylstibines, $(t Bu)$ ₂SbH has marginal volatility for MOCVD purposes (1) Torr a $24-25$ °C) and has been isolated in low vields (17%),¹¹ whereas (cyclohexyl)₂SbH could not be vacuum distilled at 1 Torr since it decomposed above 75 °C.¹² The arylstibines $(\text{Ph})_2\text{SbH},^{13}$ (Ph)SbH₂,¹⁴ and (mesityl)₂SbH¹⁵ have been synthesized in yields of **47,** 95, and **68%,** re-

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spectively. However, (Ph) ₂SbH has a low volatility (1 Torr at 118-120 °C), PhSbH₂ decomposes slowly at -70 °C under Ar with no boiling point data given, and (mesityl)₂SbH is a solid that melts at $112-115$ °C.

The objectives of this work have been to synthesize new primary and secondary stibines and to conduct a preliminary evaluation of their potential **as** MOCVD precursors. The neopentyl group $(\text{Me}_3\text{CCH}_2)$ was selected as the organic substituent in these compounds for three reasons: first, it contains no β -hydrogen atoms, thus eliminating a potentially low-energy decomposition pathway; second, the steric requirements of the substituent should hinder access to vacant coordination sites on the antimony that might be used in a migration, bimolecular, or surface-catalyzed decomposition process; third, it was anticipated that a five-carbon group should be small enough to allow a practical degree of volatility. In addition to the synthesis of these compounds, we have initiated studies to evaluate their ability to deposit InSb with Me₃In, to determine deposition parameters for a simple atmospheric reactor, and to measure gross impurities in the films.

Experimental Section:

General Techniques. The compounds described in this paper were synthesized using a combination of Schlenk and vacuum-line techniques. Although no information is available on the toxicity of these compounds, the parent compound, SbH₃, is intensely toxic, and its derivatives should be regarded likewise until data are available to indicate otherwise. All glassware used in the synthetic work was washed in **6** M HC1, **rinsed** with distilled water, and oven-dried. It was observed that not employing the acid wash resulted in the partial decomposition of the stibines as evidenced by a darkening on the surface of the glass. Other manipulations were performed in a helium-filled Vacuum Atmospheres glovebox. Diethyl ether and tetraglyme were dried over sodium/benzophenone and distilled prior to use. Deuterated benzene (Aldrich) was dried over Na/K alloy and vacuum distilled. Lithium aluminum hydride (Aldrich) was used **as** received. Antimony tribromide was prepared by the careful addition of a stoichiometric amount of bromine to antimony powder suspended in a hexane solution. 'H and 13C NMR spectra were recorded on a Bruker **MSL-300** at **300.13** and **75.468** *MHz,* respectively. 'H spectra were referenced to the residual protic peak of benzene at **7.15** ppm, while proton-decoupled 13C spectra were referenced to the center of the benzene triplet at **128.000** ppm. Vapor pressure measurements were made using a capacitance manometer. Infrared spectra were recorded **as** either gases or neat liquids between potassium bromide plates on a Perkin-Elmer **1430** spectrophotometer. Absorption intensities were determined by the method of Durkin et al.¹⁶ and are reported with the following abbreviations:
w (weak), m (medium), sh (shoulder), s (strong), v (very). Elemental analyses were performed by E&R Microanalytical Laboratories, Corona, NY.

Synthesis of (Me₃CCH₂)SbBr₂ (1). The synthesis of 1 followed a procedure published by Sowerby and Breunig for the preparation of PhSbBr₂.¹⁷ In a typical reaction, 1.00 g (2.98 mmol) of (Me3CCH2),Sb was mixed in a flask with **2.16** g **(5.96** mmol) of SbBr3 and suspended in an oil bath at **140** "C. Upon melting of the two **solids** (approximately *80-90* **"C),** the melt turned black and solidified. Continued warming to **140** "C and heating for **24** h converted the black solid to a yellowish-brown oil. Vacuum distillation of the crude product through a short-path still afforded **2.56** g **(81.0%** yield) of slightly greenish-yellow oil: bp **53-57** "C $(4.0 \times 10^{-5} \text{ Torr})$; ¹H NMR δ 0.70 (s, 9 H, SbCH₂C(CH₃)₃), 2.59 $($ s, ² **H**, SbCH₂C(CH₃)₃); ¹³C^{{1}H} δ 32.60 $($ SbCH₂C(CH₃)₃), 32.83 $(SbCH₂C(CH₃)₃), 58.24$ $(SbCH₂(CH₃)₃);$ IR (neat) 2955 (s), 2930 (sh), **2875** (m), 2860 (m), **1465 (s), 1435** (w), **1400 (vw), 1380** (m), **1365 (s), 1265** (w), **1235 (s), 1130** (w), **1095 (vw), 1020 (vw), lo00** **(vw), 740** (w), **605** (w) cm-'. Elemental Anal. Celcd for C5H11SbBr2: C, **17.03;** H, **3.14;** Br, **45.31.** Found C, **17.19;** H, **3.07;** Br, **45.66.**

Synthesis of (Me₃CCH₂)SbH₂ (2). Although several different reaction schemes were attempted in order to prepare **2,** only one will be described in depth based on its superiority to the others. These deficient schemes will be discussed later in the paper. To **0.112** g **(2.95** mmol) of lithium aluminum hydride suspended in **8** mL of tetraglyme stirred at **-25** "C, **1.07** g **(3.03** mmol) of (Me3CCH2)SbBr2 in **8** mL of tetraglyme at 0 "C was added dropwise via cannula with stirring. Immediately upon introduction of the dibromide to the lithium aluminum hydride solution, a darkening was observed with a small evolution of gas. Addition of the dibromide solution was made over a period of **5-10** min. Following addition, the reaction flask was suspended in an ice bath and evacuated on a high-vacuum line. Volatile materials from the reaction were trapped and fractionated by passing through a **-78** "C trap and a **-196** "C trap. The contents of the **-78** "C trap were examined and found to be a white solid that melted with slight warming (approximately **-50** "C). Transfer of the volatile liquid to a tared container afforded **340** mg **(57.6%** vield) of liquid identified as $2.$ ¹H NMR δ 0.86 **(s, 9 H**, $SbCH_2C(CH_3)_3$, 1.62 (t, ³J = 4.8 Hz, 2 H, $SbCH_2C(CH_3)_3$), 1.70 $(t, {}^{3}J = 4.8 \text{ Hz}, 2 \text{ H}, \text{SbH}_2)$; ¹³C¹H) δ 21.04 (SbCH₂C(CH₃)₃), 31.34 $(SbCH₂C(CH₃)₃$, 31.70 $(SbCH₂C(CH₃)₃$; IR (gas) 2965 (vs), 2915 (sh), **2875** (w), **1870** (vs), **1470** (m), **1365** (m), **1240** (m), **1140** (m), **810** (m), **790** (w), **645** (w), **575** (w). Elemental Anal. Calcd for Sb, **62.52.** CSH&3b C, **30.81;** H, **6.72;** Sb, **62.47.** Found: C, **31.08;** H, **6.49;**

Synthesis of $Me₃CCH₂$ **₂SbH** (3). The preparation of the secondary stibine 3 was accomplished by the reduction of the previously reported (Me3CCH2)2SbI.1s To **0.065** g **(1.6** mmol) of LAH suspended in 5 **mL** of diethyl ether at **-50** OC, **1.75** g **(4.48** mmol) of $(\text{Me}_3\text{CCH}_2)_2$ SbI in 10 mL of diethyl ether was added slowly with stirring. The mixture was then held at -50 °C for **2** h, followed by gradual warming to ambient temperature. Following stirring overnight, the diethyl ether was removed in vacuo at -20 °C (0.01 Torr) to leave an oily residue. This residue was transferred to a short-path still by cannula and vacuum distilled to afford 0.673 g (56.8% yield) of a colorless liquid: bp **11-12 °C** (0.01 Torr); ¹H NMR δ 0.98 (s, 18 H, $SbCH_2C(CH_3)_3$), **1.28** (m, 2 H, $^{2}J = 12.0$, $^{3}J = 5.5$ Hz, SbCH₂C(CH₃)₃), 2.03 (m, **2 H,** $^{2}J = 12.0$, $^{3}J = 5.5$ Hz, SbCH₂C(CH₃)₃), 2.13 (quintet, 1 H, $H_2C(CH_3)_3$, 32.90 $(SbCH_2C(CH_3)_3)$; IR (neat) 2955 (vs), 2940 (sh), **2900** (m), **2870** (m), **1840 (s), 1470 (s), 1385** (m), **1365** (m), **1240 (s), 1140** (sh), **1130** (m), **1100** (w), **1020** (w), **1005** (w), **790** (m), **755** (w), **740** (w), **650** (w), 555 (w). Elemental Anal. Calcd for C1&&b: C, **45.32;** H, **8.75;** Sb, **45.94.** Found: C, **45.14;** H, **8.87;** Sb, **46.08.** $^{3}J = 5.5$ Hz, SbH); ¹³C(¹H) δ 28.72 (SbCH₂C(CH₃)₃), 32.09 (SbC-

Chemical Vapor Deposition Experiments. Silicon or gallium arsenide **(100)** wafers with a thickness of **0.4** mm were cut into pieces ranging from 10 to 20 mm/side. The pieces were then degreased with detergent, washed with distilled water, isopropyl alcohol, **boiling** trichloroethylene, and isopropyl alcohol, and finally blown *dry* with a stream of nitrogen. The piece of cleaned wafer (polished side up) was then placed on a static, graphite susceptor within an atmospheric pressure, vertical growth, Pyrex reactor. A smaller piece of substrate (polished side down) was placed on top of the clean wafer to serve as a mask. The vertical reactor contained three separate inlets arranged in a triangular fashion with spacing between the inlets of approximately one inch. For all CVD growth runs, the graphite susceptor was positioned about **2.5** in. below the inlets and inductively heated with an rf source operating between **250** and **800** kHz to calibrated temperatures of **350-450** "C depending on conditions of the growth run. Using Cajon Ultra-Torr fittings and stainless steel flexible tubing, two glass bubblers (one containing trimethylindium, Alfa, electronic grade, and the other containing a neopentylantimony source) were connected to two of the reactor inlets. The third inlet was connected to a hydrogen flow source (push gas). Flows through both bubblers and the push gas were controlled with calibrated flowmeters. The hydrcgen carrier and push gas were reagent grade

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and passed through a molecular sieve cartridge. **Gas** flows for the two antimony sources, trimethylindium, and the push **gas** were in the ranges **50-150** standard cm3/min (sccm), **25-50** sccm, and **0.5-2.0 L/min,** respectively, depending on conditions of the specific growth run. During growth runs both trimethylindium and $(Me₃CCH₂)₂SbH$ were maintained at ambient temperature (ca. 23 $^{\circ}$ C), while (Me₃CCH₂)SbH₂ was held at 0 $^{\circ}$ C. Following the termination of a growth run, the susceptor was heated for an additional **3-5** min with purging after stopping the source flows; purging was then continued **as** the system was allowed to cool to ambient temperature. The substrate was subsequently removed from the reactor and characterized. A **total** of eight individual growth runs were made for each antimony source $\text{(Me}_3\text{CCH}_2)_2\text{SbH}$ and $(\text{Me}_3\text{CCH}_2)\text{SbH}_2$.

Characterization of Thin Films. The thin films grown by the chemical vapor deposition experiments were characterized by Auger electron spectroscopy (AES) and X-ray diffraction (XRD). The Auger spectra were obtained on a Physical Elec**tronics** Industries Model **10-155** equipped with a **cylindrical** mirror analyzer. Spectra of as-mounted films were recorded, followed by sputtering with Ar⁺ ions for 2 min at a pressure of 5×10^{-5} Torr with a beam energy of **700** eV and **a** beam current of **16** MA to remove carbon and oxygen contaminants from the surface of the **film.** Data from the sputtered samples are reported here. A reference **spectrum** for InSb was obtained on a piece of molecular beam epitaxy (MBE) grown InSb (bulk mobility 2×10^5 cm²/V **s)** and gave an In/Sb ratio of **1.13.** X-ray diffraction patterns were obtained on Philips 12045 θ -20 diffractometer equipped with Cu K_{α} radiation. Thickness measurements were obtained with a Sloan Dektak **900051** profilometer.

Results and Discussion:

Both (Me₃CCH₂)SbH₂ and (Me₃CCH₂)₂SbH were synthesized by reduction of the corresponding organoantimony halide species with lithium aluminum hydride. The two stibines have been prepared and isolated in **57-58%** yield (reaction des involving up to **3** g of product or approximately $10-15$ mmol). A scheme showing the reaction is depicted in eq **1.** Although eq **1** describes the

$$
(Me3CCH2)ySbX3-y + LiAlH4 \xrightarrow[low temperature]{\text{low temperature}}
$$

\n
$$
(Me3CCH2)ySbH3-y
$$

\nfor $y = 2$, $X = I$
\nfor $y = 1$, $X = Br$

reduction of only the bromide and iodide species, it is worth noting that we attempted the reduction of both **dineopentylchlorostibine** and neopentyldichlorostibine. However, in our hands these compounds produced little or no isolable Sb-H containing products. Thus, it would appear that the nature of the halide is crucial for optimizing the yield of the antimony hydride complex. Also, we have found no dependence on the stoichiometry of $LiAlH₄$ used (0.25-1.0 equiv of LiAlH₄/equiv of haloantimony compound) with respect to the yield of Sb-H compound obtained.

Both stibines are colorless, volatile liquids at room temperature and are soluble in hydrocarbons and ethers. The vapor pressure of (Me3CCH2)2SbH is **0.5** Torr at **24** "C, whereas that of the substantially more volatile $(Me_3CCH_2)SbH_2$ is 5.5 Torr at 0 °C. In addition to these physical properties, the two liquids turn to white solids when exposed to air but are nonpyrophoric in small quantities. Both compounds exhibit surprising stability at ambient temperatures in the absence of light **as** evidenced by the primary stibine showing no darkening or formation of noncondensable gas when stored under vacuum for 2 months and the secondary stibine remaining unchanged over a period of 3 months when stored under an argon atmosphere. However, we have observed a

Figure 1. Infrared spectra of $(Me_3CCH_2)SbH_2$ and $(M_{\rm{e}_3}CCH_2)_2SbH$ showing the Sb-H region.

Figure 2. ¹H NMR spectrum of $Me₃CCH₂$ ₂SbH.

blackening of the walls of a tube containing either of the hydrides when exposed to fluorescent light for a period of at least 2 weeks. In the case of $(Me_3CCH_2)_2SbH$, we have determined a photodecomposition product to be the distibine, $(Me_3CCH_2)_4Sb_2$ (confirmed by NMR and X-ray crystallography).¹⁹ We believe this product to arise from the evolution of H_2 along with the coupling of two $(Me_3CCH_2)_2Sb-$ fragments.

In addition to their physical properties, the two stibines have been characterized by infrared and NMR spectroscopy. The infrared spectra of the two hydrides are dominated by a very intense absorption associated with the Sb-H stretch. For the neat liquid $(Me_3CCH_2)_2SbH$, the Sb-H stretch is observed at 1840 cm^{-I}. In the case of $(Me₃CCH₂)SbH₂$ in the gas phase, the band is centered at 1870 cm-'. The two Sb-H regions are shown in Figure **1.** These values compare favorably with $SbH₃$, which exhibits a central Sb-H stretch at 1890 cm^{-120} and is some 230 cm^{-1} lower in energy than the corresponding neopentylarsines.²¹ The ¹H NMR spectrum for $(Me_3CCH_2)_2SbH$ in Figure 2 shows a singlet for the methyl groups while the methylene protons and antimony hydrogen form an AA'BBX system. The two nonequivalent methylene hydrogens occur at 1.28 and 2.03 ppm (showing ${}^2J_{AB} = 12.0$ Hz and ${}^3J_{AX} = {}^3J_{BX}$ $= 5.5$ Hz) and are due to the different magnetic environments imposed by the pyramidal structure and the reduction in symmetry resulting from having only two neopentyl groups bound to the antimony center. The antimony hydrogen resonance is identified by the quintet pattern at 2.13 ppm with a coupling of 5.5 Hz. A ¹³C $\{$ ¹H $\}$ spectrum from $(Me₃CCH₂)₂SbH$ shows the expected three

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Table I. XRD Data for InSb Film Grown Using $(Me_3CCH_2)_2SbH$

Figure 3. Auger spectrum of InSb film **employing trimethyl**indium and $(\text{Me}_3\text{CCH}_2)_2\text{SbH}$ as precursors.

resonances with the methylene $(-CH₂)$ carbon appearing at **28.72** ppm. This chemical shift for the methylene carbon is some 9.10 ppm upfield from $(M_{\rm{e_3}}CCH_2)_3Sb^{18}$ and can be attributed to the shielding nature of the hydrogen atom. The ¹H spectrum for $(Me₃CCH₂)SbH₂$ is much simpler than that observed for $(Me₃CCH₂)₂SBH$ in that both the methylene hydrogens and the antimony hydrogens appear **as** triplets at **1.62** and **1.70** ppm respectively. The ${}^{13}C_1{}^{1}H$ spectrum for $Me_3CCH_2)SbH_2$ also shows three resonances. However, the methylene carbon exhibits a further upfield **shift** of **7** ppm from that of the methylene carbon in (Me3CCH2)2SbH with a chemical shift of **21.04** ppm.

Our ultimate interest in these compounds has been to explore their potential utility **as** precursors for use in MOCVD. Specifically, we have carried out preliminary investigations of both stibines in conjunction with trimethylindium in attempts to grow indium antimonide **films** in a very simple reactor. Our goal has been to demonstrate growth of indium antimonide films and to identify whether these compounds had potential **as** an alternative antimony source compound.

Despite the low vapor pressure of $(Me_3CCH_2)_2SbH$ at ambient temperature (0.5 Torr), we were able to obtain films of polycrystalline indium antimonide on a silicon **(100)** substrate. Typical growth conditions employed a deposition temperature of **400** "C and an approximate V/IIl ratio of **3.5.** The polycrystalline nature of the indium antimonide was confimed by X-ray diffraction patterns and showed no preferred orientation.²² A listing of observed **28** values, *d* values, materials, and crystallographic orientation appears in Table I. In addition to the peaks associated with indium antimonide, we **also** observed **peaks** for antimony and/or indium. We attribute this to the simple design of our reactor, which does not allow for fine control of deposition parameters. Examination of the films by Auger electron spectroscopy showed various ratios of indium to antimony. Figure 3 shows an Auger spectrum

Table 11. XRD Data for InSb Film Grown Using

d_{obsd} , Å $d_{\text{lit.}}$, Å identity $2\theta_{\rm obad}$ 23.70 InSb 3.751 3.74 Sb^b 28.70 3.108 3.109 102 (31) 39.30 2.291 2.290 InSb 220 (34) Sb 39.95 2.255 2.248 014 (9) 41.95 Sb 2.152 2.152 110 (6) 46.40 InSb 1.955 1.953 311 (18) 48.20 1.886 Sb 1.878 006 (14) 51.60 1.770 1.770 Sb 022(5) 56.75 1.621 InSb 1.620 400 (2) 59.40 Sb 1.555 1.555 204 (1) 62.50 1.485 InSb 1.486 331 (3) 4 See ref 22. $\sqrt[3]{2}$ See ref 23.	$(Me3 CCH2) SbH2$						
		orientation $(I/I_0, \mathcal{D})$					
		111 (100)					
Sb			404/410 e V				
In 460/468 eV							

I 300 **eV** 400 **eV 500 eV 600 eV** 200 **eV Figure 4. Auger spectrum of InSb** film **employing trimethyl**indium and $(\text{Me}_3\text{CCH}_2)\text{SbH}_2$ as precursors.

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for our best film obtained at the previously stated growth conditions. Central to the spectra are peaks at **402/409** eV for indium and $456/464$ eV for antimony. In this spectrum, the In/Sb ratio of **1.08** compares favorably **to 1.13** from a sample of MBE-grown indium antimonide analyzed in our spectrometer. A small peak at **277** eV indicates carbon contamination in this film to be a minimum of **1%** , the approximate detection level of Auger spectroscopy. Nomarski photographs of these films exhibited a poor morphology with large grains (approximately $1-\mu m$ diameter) and numerous defects. This may be attributed to the nature of our reactor, incorrect growth conditions, or a **20%** lattice mismatch between silicon and indium antimonide. Profilometry results indicated a low growth rate that varied from 0.1 to 0.3 μ m/h.

Similarly, we have employed $(Me_3CCH_2)SbH_2$ in attempts to grow films of indium antimonide. The primary stibine was of particular interest to us due to its enhanced volatility in comparison to $(Me₃CCH₂)₂SbH.$ In our growth **runs,** we were able to obtain the deposition of polycrystalline indium antimonide on either a silicon or gallium arsenide (100) substrate at 350 °C. A typical run employed a V/III ratio of approximately **5.5-6.** The polycrystalline nature of the film was established by XRD with no preferred orientation exhibited (data given in Table 11). Profilometry resulta indicated an increase in growth rate, as compared to $(Me₃CCH₂)₂SbH$, with values ranging from 0.4-0.6 μ m/h. Examination of the films by Auger spectroscopy showed an In/Sb ratio that varied from **0.64** to **1.40** depending on growth conditions. However, the Auger spectra contained no evidence of carbon in the 270-eV region indicating that the amount of carbon present was below the Auger detection limit of approximately **1%** (see

⁽²²⁾ Joint Committee on Powder Diffraction Standards (JCPDS), File

Figure 4). No attempts were made to quantify carbon contamination using more sensitive techniques since the presence of excess antimony in the film and the poor morphology (granular in nature), **as** seen from visual inspection and **Nomarski** micrographs, indicated we had not optimized the deposition conditions. Nevertheless, the absence of carbon at the percent level in these films, **to**gether with the stability and volatility of the primary stibine, suggests it is a suitable candidate for further growth studies under more controlled conditions.

Conclusions:

From this study we have found that stable, volatile primary and secondary antimony hydrides can be **syn**thesized in reasonably good yields. These compounds are significantly more stable and/or volatile than other organoantimony hydrides reported in the literature. This may be attributed to the presence of large organic substituents with no β -hydrogens. Our studies have also shown that **these** hydrides are transportable and *can* be used to deposit InSb. In particular, these compounds can be used with trimethylindium to deposit InSb at moderately low temperatures for organostibines (400 $^{\circ}$ C for (Me₃CCH₂)₂SbH and 350 °C for $(Me₃CCH₂)SbH₂)$. The primary stibine is extremely attractive because of its high volatility and low decomposition temperature, characteristics of interest in growing ternary antimonides or metastable antimonidebased materials.

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 $(Me_3\text{CCH}_2)_3\text{Sb}$, 99715-60-3; SbBr_3 , 7789-61-9; $(Me_3\text{CCH}_2)_2\text{SbI}$, 141612-23-9; Me₃In, 3385-78-2; InSb, 1312-41-0. **Registry NO.** 1,129135-92-8; 2,129135-94-0; 3,129135-93-9;

Synthesis and Evaluation of Copolymers of *(tert* **-Butoxycarbonyloxy)styrene and Chemically Amplified Deep-UV Imaging Materials (2 -Nitro benzyl) styrene Sulfonates: Single- Component**

J. E. Hanson,**t E. Reichmanis,* F. M. Houlihan, **and** T. X. Neenan

AT&T Bell Laboratories, Murray Hill, New Jersey 07974

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Copolymers of **(tert-butoxycarbony1oxy)styrene** (t-BOC-styrene) and (2-nitrobenzy1)styrene sulfonates were prepared to give hydrophobic, photosensitive materials. Exposure to ultraviolet light converted the pendant 2-nitrobenzyl sulfonate esters to sulfonic acid moieties, and this photogenerated acid catalyzed the removal of the t -BOC groups on subsequent heating. The resulting material was a copolymer of 4hydroxystyrene and styrenesulfonic acid, which was hydrophilic and soluble in aqueous base. Copolymers with different compositions (monomer ratios) and molecular weights were synthesized, and the resulting materials were **analyzed** by thermal and spectroscopic methods and evaluated for lithographic performance **as** positive tone resists. *As* the percentage of the photosensitive 2-nitrobenzylstyrene sulfonate monomer was increased, the lithographic sensitivity improved but the thermal stability of the copolymers decreased. The photoresponse of some of the materials was inhomogeneous, apparently the result of inhomogeneous distribution of acid-generating monomers in the polymers.

The needs of the semiconductor industry continue to give impetus to the investigation of new chemistries for photoresists. These needs include the resolution of smaller features, improved sensitivity (or photospeed), and greater process latitude (the ability to tolerate more variation in processing conditions).' Chemically amplified deep UV photoresists are one promising class of materials for patterning in the submicron regime $(0.25-1.0 \ \mu m)^2$ Deep-UV is defined **as** light in the wavelength range **220-280** nm, though typically 248-nm KrF excimer laser radiation is employed. Chemical amplification refers to the increase in sensitivity achieved by harnessing the effect of each absorbed photon to catalytic resist chemistry. 3 A typical chemically amplified resist contains two components: a matrix polymer and a photoactive additive. Absorption of light by the photoactive compound generates the catalytic agent, usually a strong acid. Subsequent heating

Introduction drives an acid-catalvzed chemical reaction of the matrix polymer, generally a deprotection or cross-linking reaction. The end result is different solubility properties for the exposed and unexposed areas of the resist film, allowing formation of the desired image.

A chemically amplified resist we have studied previously consists of poly[**(tert-butoxycarbony1oxy)styrene-co-sulfur** dioxide] (PTBSS) and a photoacid generator.^{3d} We have

Current address: Department of Chemistry, Seton Hall University, South **Orange,** NJ 07079.

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